

# Using Quasars to Measure the Earth

Bill Petrachenko, Nov 5, 2007, University of New Brunswick

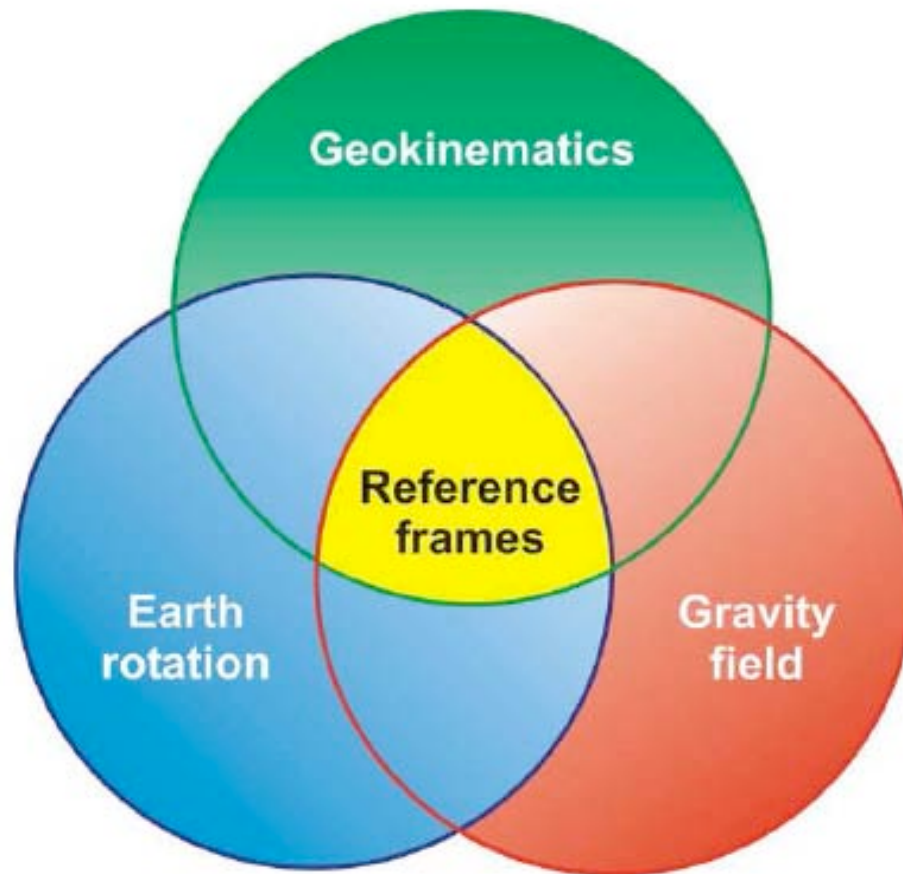


Natural Resources  
Canada

Ressources naturelles  
Canada

Canada 

# The Three Pillars of Geodesy



# What is a geodetic reference system?



- It is an idea, a set of rules used to assign coordinates to points
- In general, it includes:
  - Location of an origin
  - Orientation of axes
  - Definition of scale
- In practice, it also requires:
  - Definition of physical constants
  - Definition of physical models
- A simple example is a 2-d Cartesian system.

# What is a geodetic reference frame?

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- It is a realization, in other words, it is the way we gain *access* to a reference *system*
- In practice, it is the coordinates of a set of points determined from specific observations analyzed with appropriate models.

# Historical Progression



- Historically, reference systems progressed from:
  - local -> regional -> national -> continental
- Since classical surveying requires line of sight measurements the continents couldn't be connected
- As a result, each continent eventually had a separate reference frame with different:
  - Ellipticity, origin, and orientation
- The oceans could only be bridged using less precise positional astronomy

# Launch of Sputnik I, the Dawn of the Space Age



- On Oct 4 1957, the USSR launched the first ever man-made object into orbit, Sputnik I.
- Although not intended for geodesy, the new orbital capability allowed the continents to be connected through the mutual observation of satellites
- This heralded the era of *space geodesy*, and as a result, globally consistent geodetic reference systems could now be envisioned!
- Geos-A was launched in 1965.

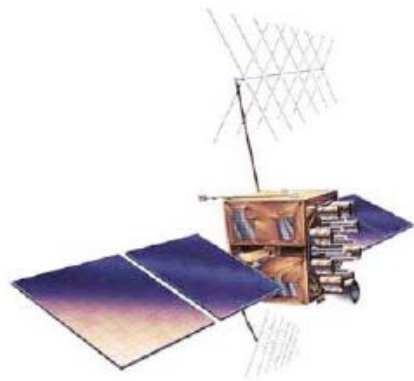
# Space Geodesy Techniques

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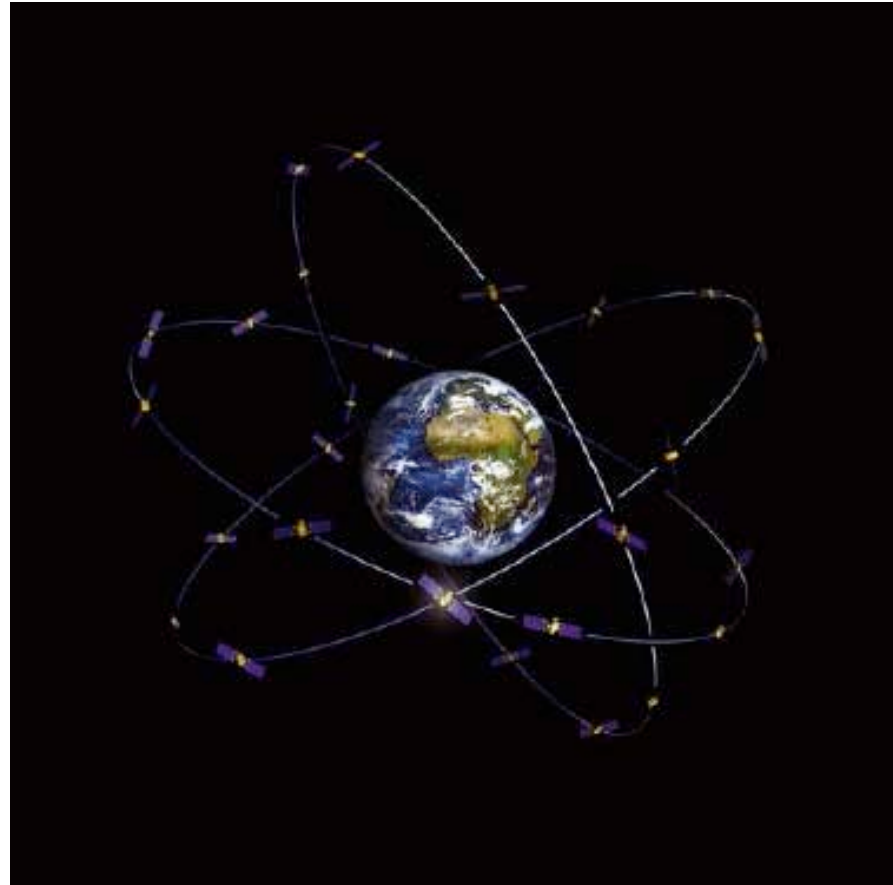


- The four main space geodesy measurement techniques (not including those related to gravity) are:
  - Global Navigation Satellite Systems (GNSS)
  - Satellite and Lunar Laser Ranging (SLR and LLR)
  - Very Long Baseline Radio Interferometry (VLBI)
  - DORIS

# Microwave Satellite Systems (GNSS)



- TRANET
- GPS
- GLONASS
- GALILEO
- DORIS





# Laser Ranging



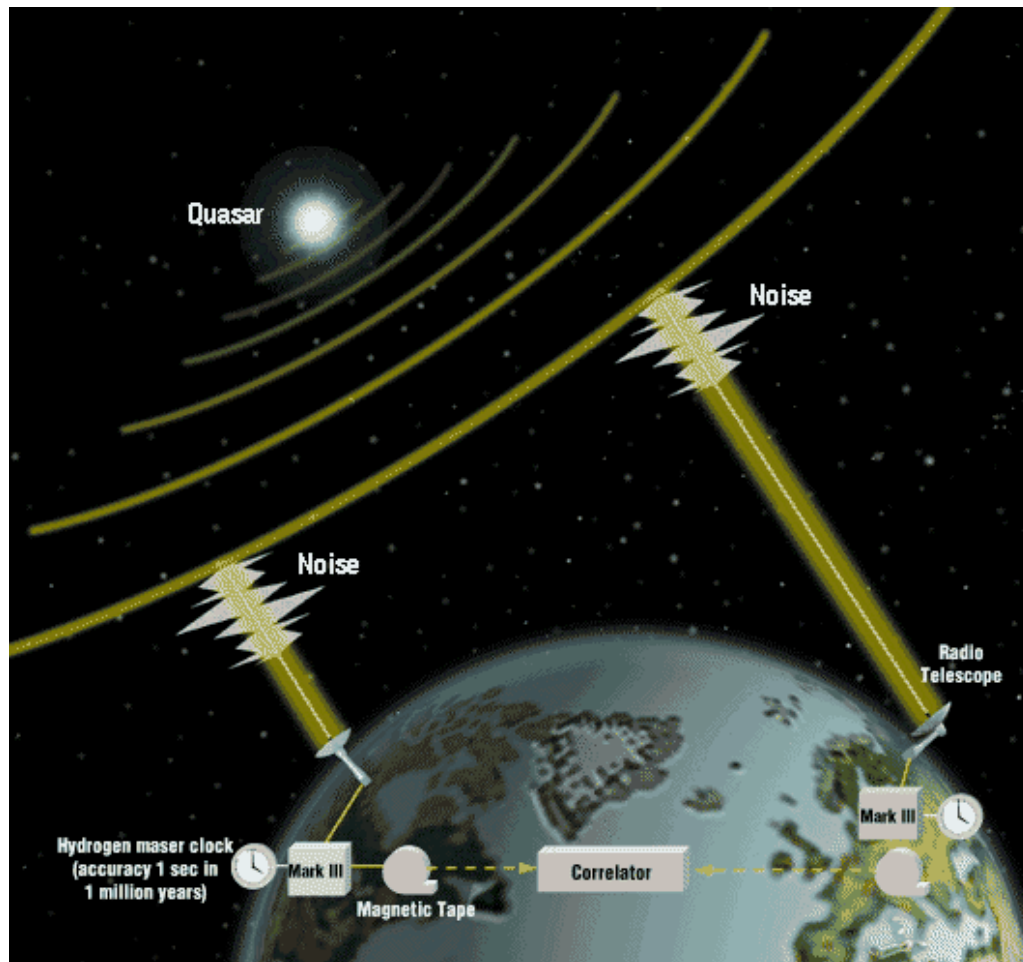
- Satellite Laser Ranging (1976)
- Lunar Laser Ranging (1969)



# Very Long Baseline Interferometry (VLBI)



- Invented by Canadian Astronomers in 1967
- Required the invention of:
  - Atomic clocks
  - Broadband tape recorders.



# Reference Frames



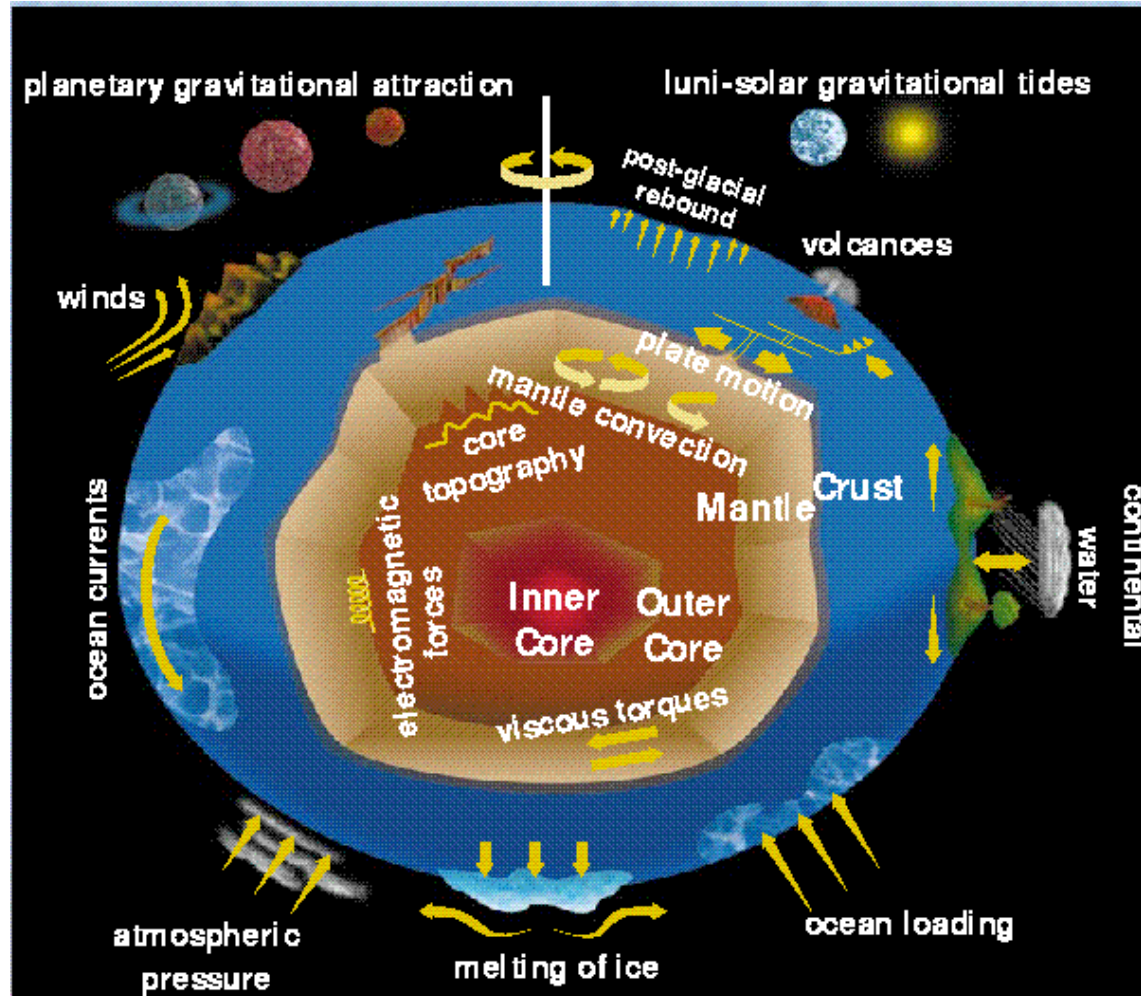
- On the global level, space geodesy (not including gravity) has three main components:
  - International Terrestrial Reference Frame (ITRF)
  - International Celestial Reference Frame (ICRF)
  - Earth Orientation Parameters (EOP)

# International Terrestrial Reference System (ITRS)

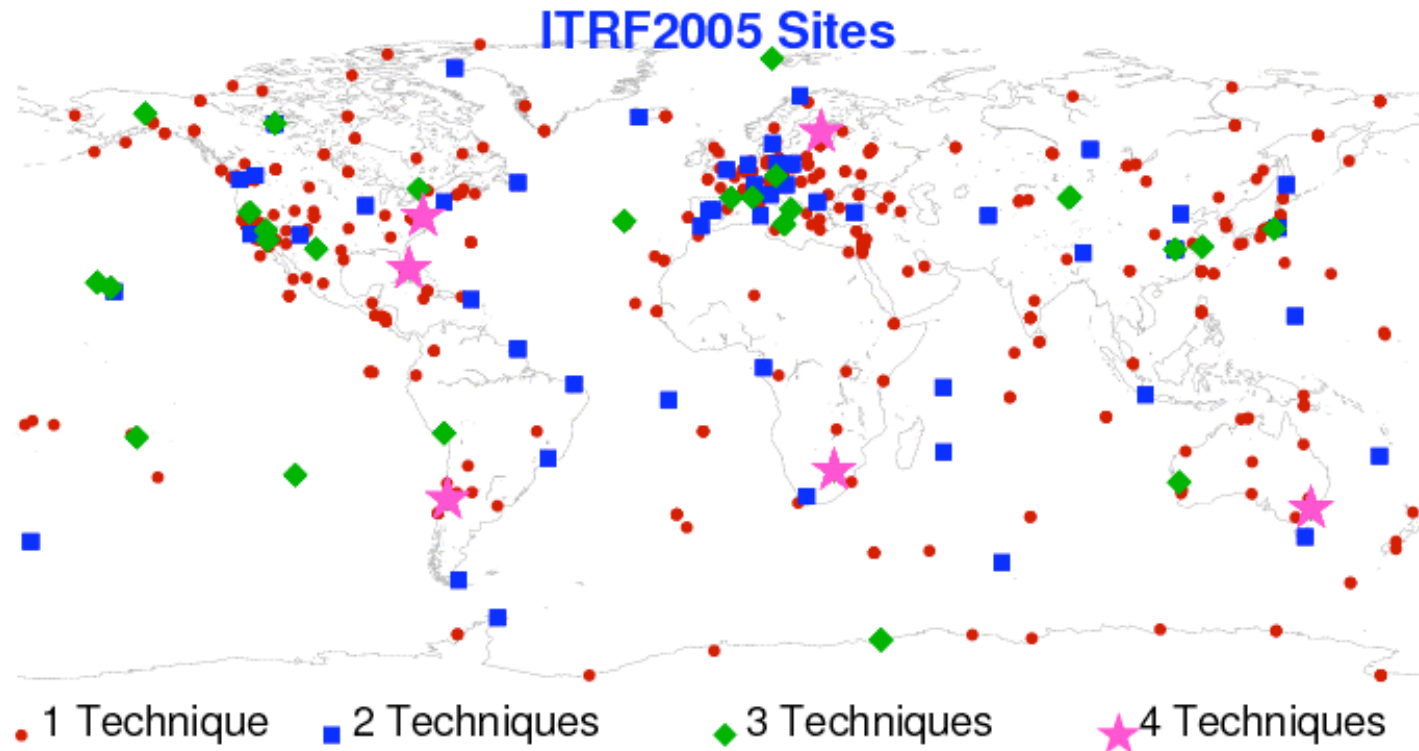


- Origin -> Centre of mass of the whole earth
  - Determined by satellite systems only, mainly SLR
- Orientation -> Equatorial system, given by the BIH orientation at 1984.0
- Scale -> Unit of length, metres
  - Determined by VLBI, SLR?
- Time evolution
  - No-net-rotation
  - Geophysical models

# Restless Planet



# International Terrestrial Reference Frame (ITRF) – coordinates and rates



Precision of best sites: 1-3 mm position; 0.1-0.3 mm/yr  
Most recent ITRF produced in 2005



# International Celestial Reference System (ICRS)

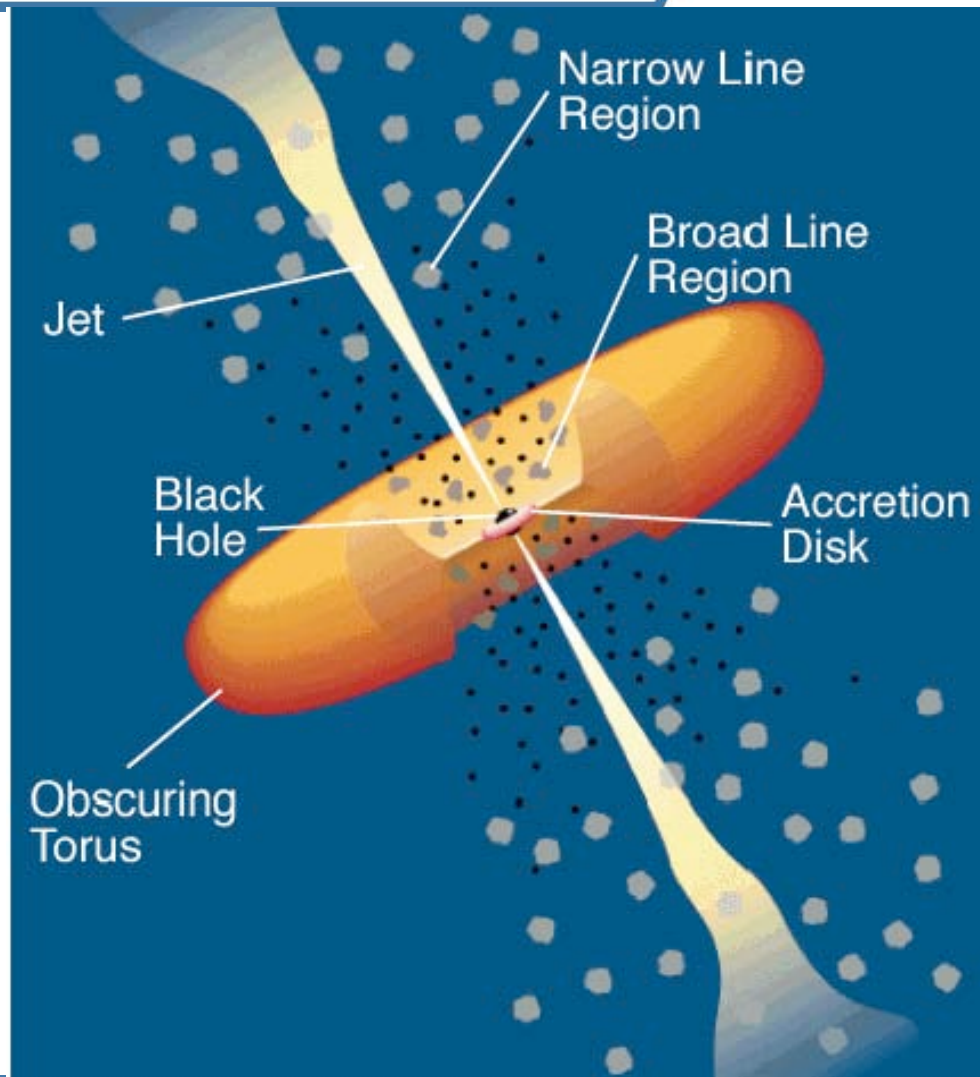


- Origin -> Solar System Barycentre
- Orientation -> Equatorial system, given by
  - Mean equator of J2000.0
  - Origin at dynamical equinox of J2000.0
- Time evolution?
  - Required regular monitoring

# Source Jet Model

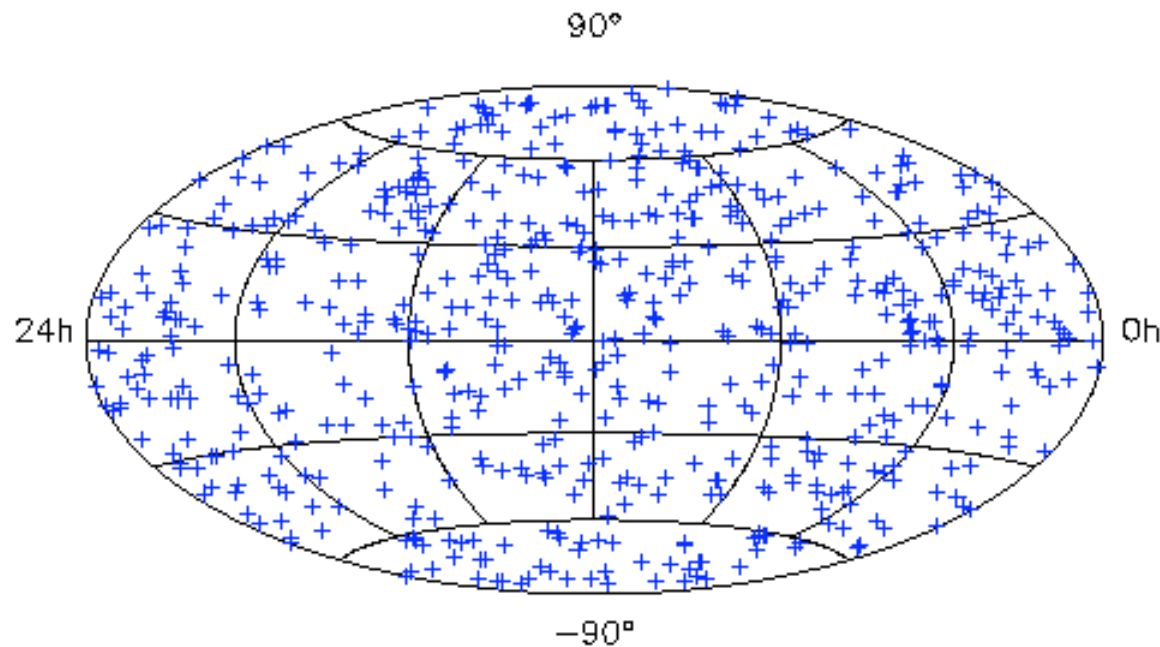


- Positionally stable point is the dense Black hole at the core
- Only the jets are visible to VLBI
- Unfortunately for geodesy, the jets are dynamic.





# International Celestial Reference Frame (ICRF)

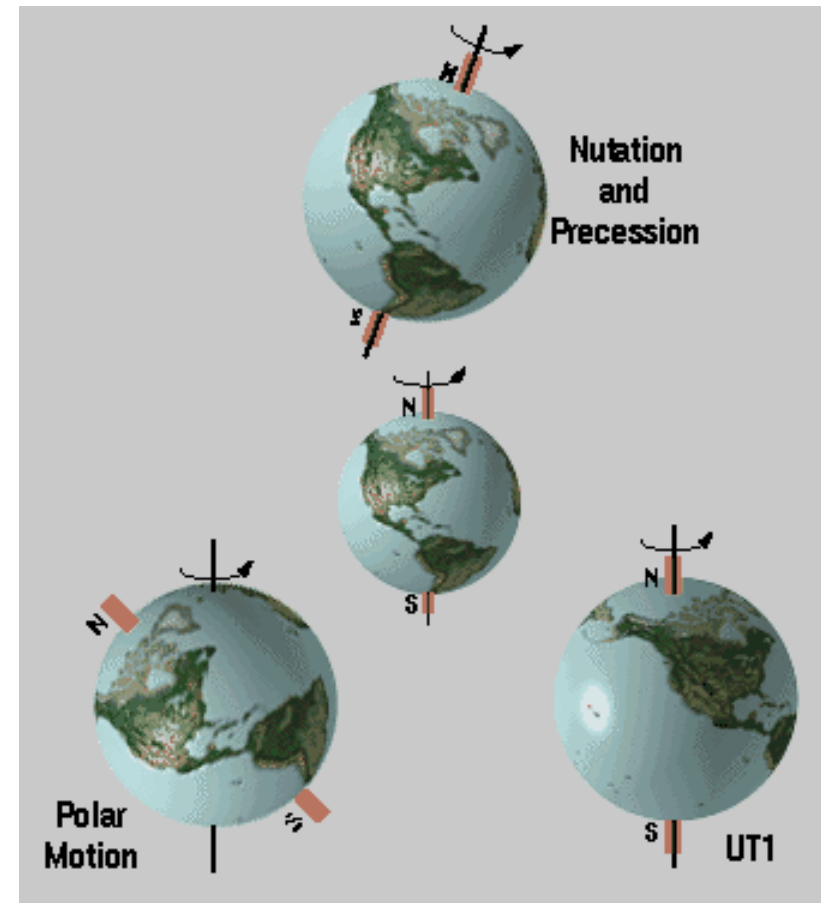


- Over 700 sources, 212 of them are defining sources
- New ICRF expected in 2009

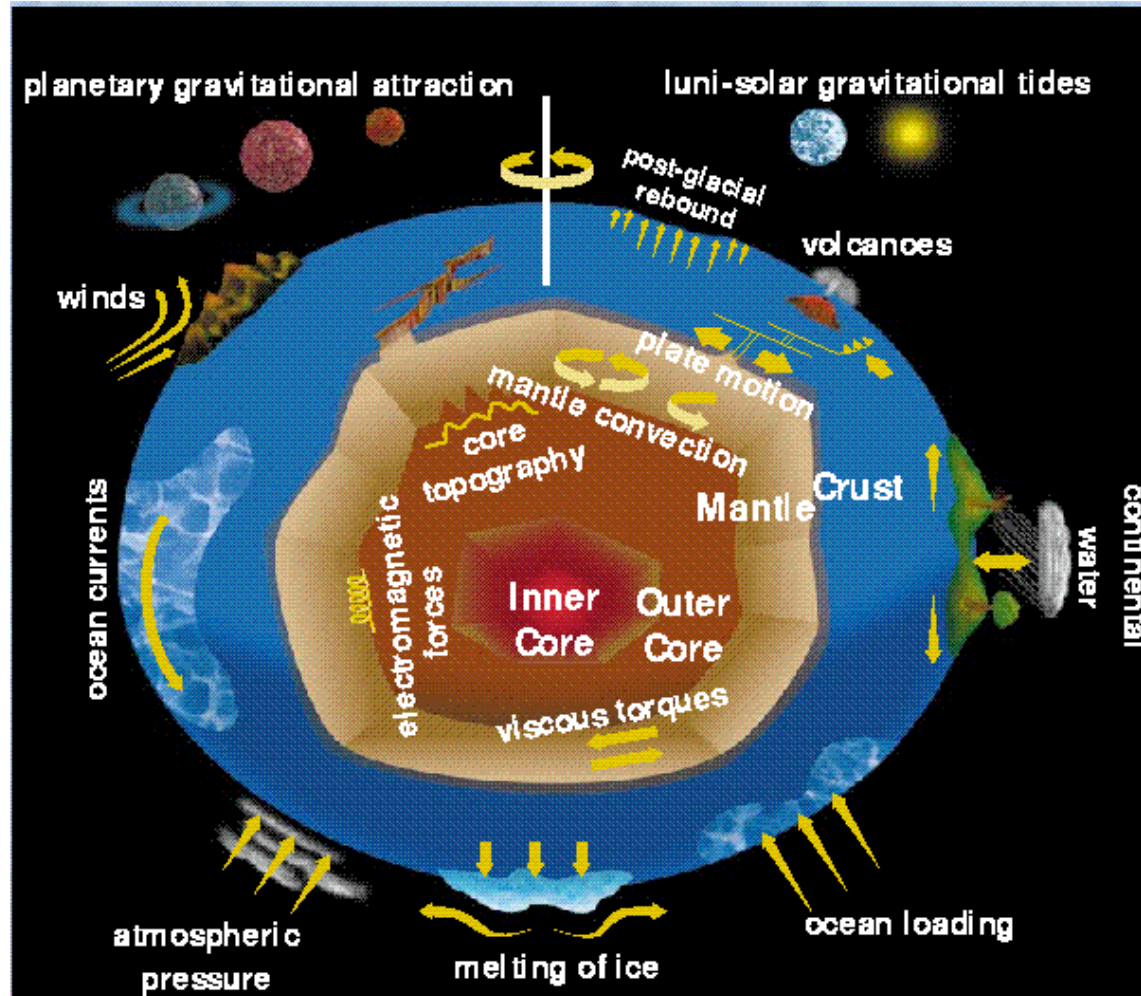
# Earth Orientation Parameters



- Why do we care?
  - Dynamical equations expressed in a non-rotating frame
- UT1, rotation angle
  - VLBI only
- Precession and nutation, orientation in space
  - VLBI
- Pole position, geographic location
  - All techniques



# Restless Planet



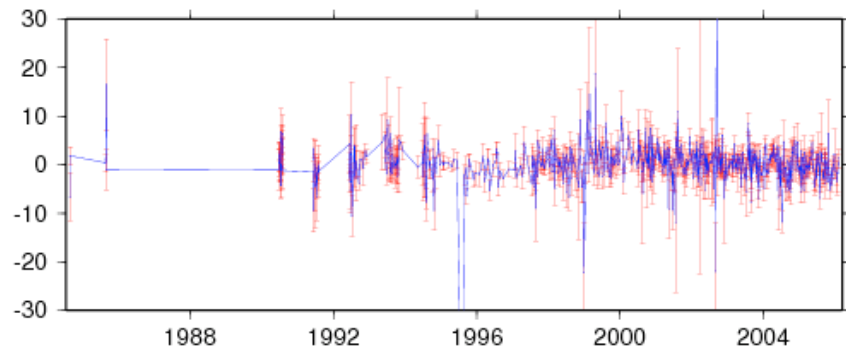
# What is VLBI's role in space geodesy?



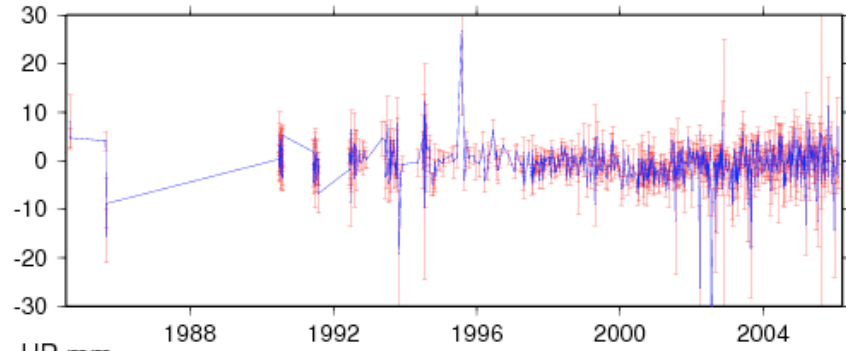
- Definition of the Celestial Reference Frame (ICRF)
  - 212 Quasars
- Determines all Earth Orientation Parameters (EOP)
  - Unique for UT1 and nutation
- Definition of the Terrestrial Reference Frame
  - Especially Scale

### 40104S001 7282 Residuals

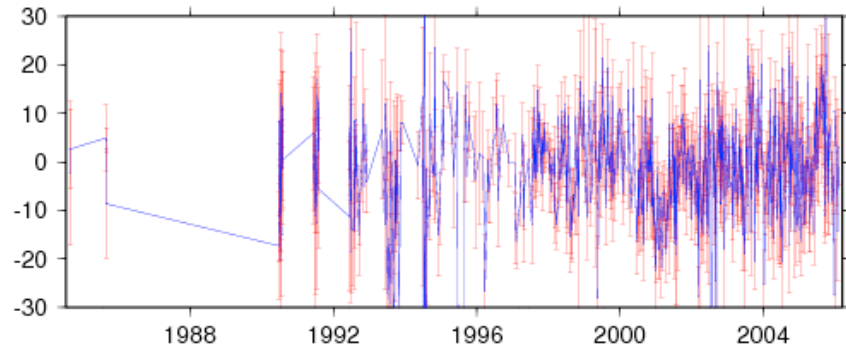
EAST mm



North mm

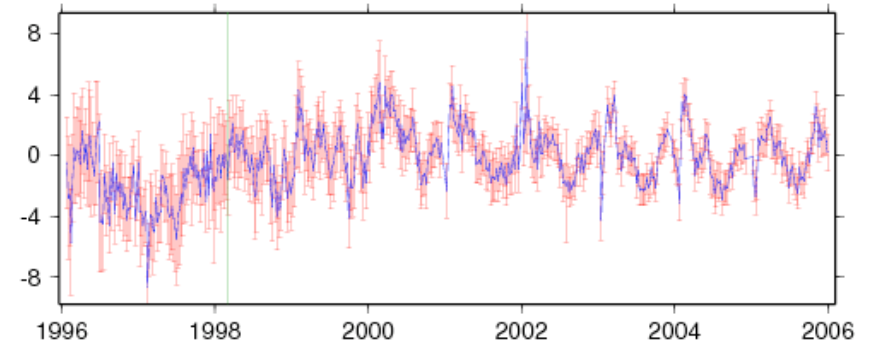


UP mm

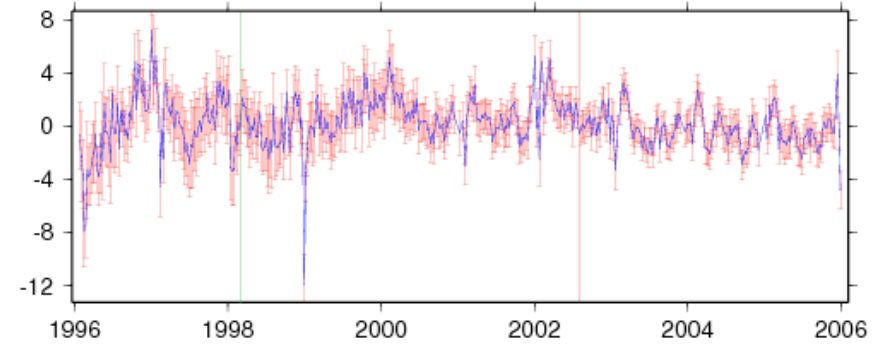


### 40104M002 ALGO Residuals

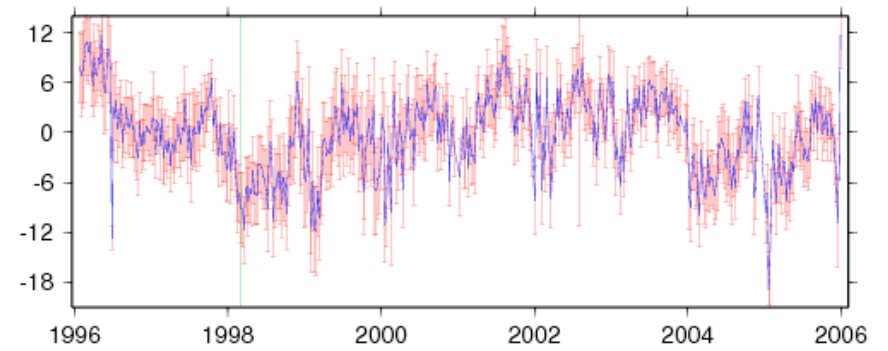
EAST mm



North mm



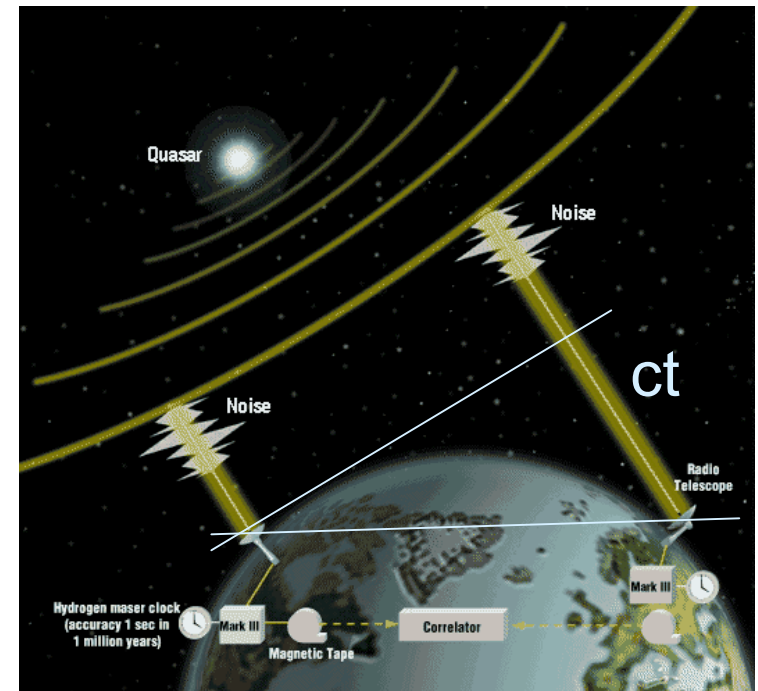
UP mm



# How does VLBI work?



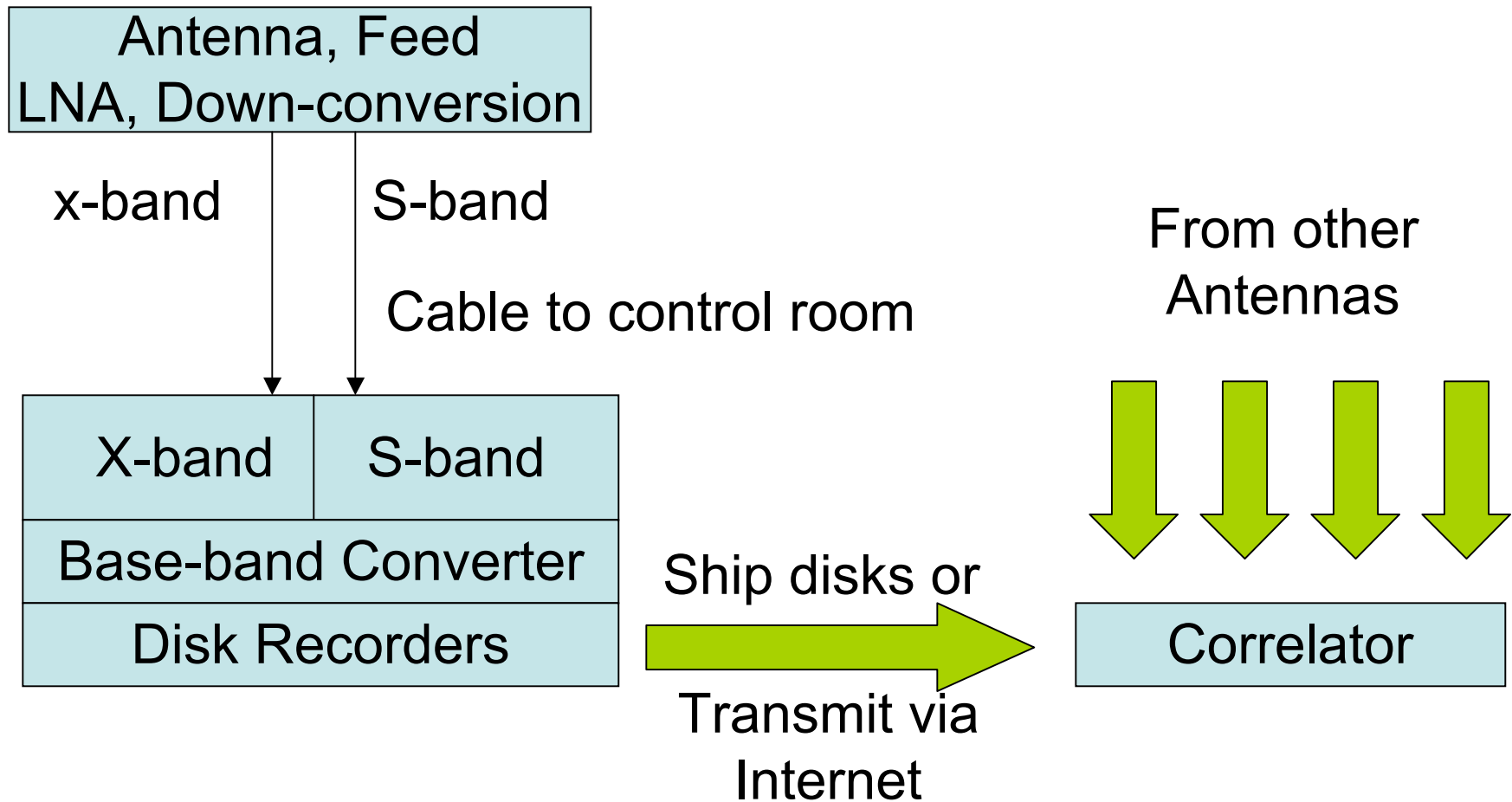
- The delay between the time of arrival of a signal at two antennas is measured
- Using the speed of light,  $c$ , this can be interpreted as a distance.
- That distance is the component of the baseline in the direction of the source
- If many sources are observed, the full 3-D vector baseline can be determined



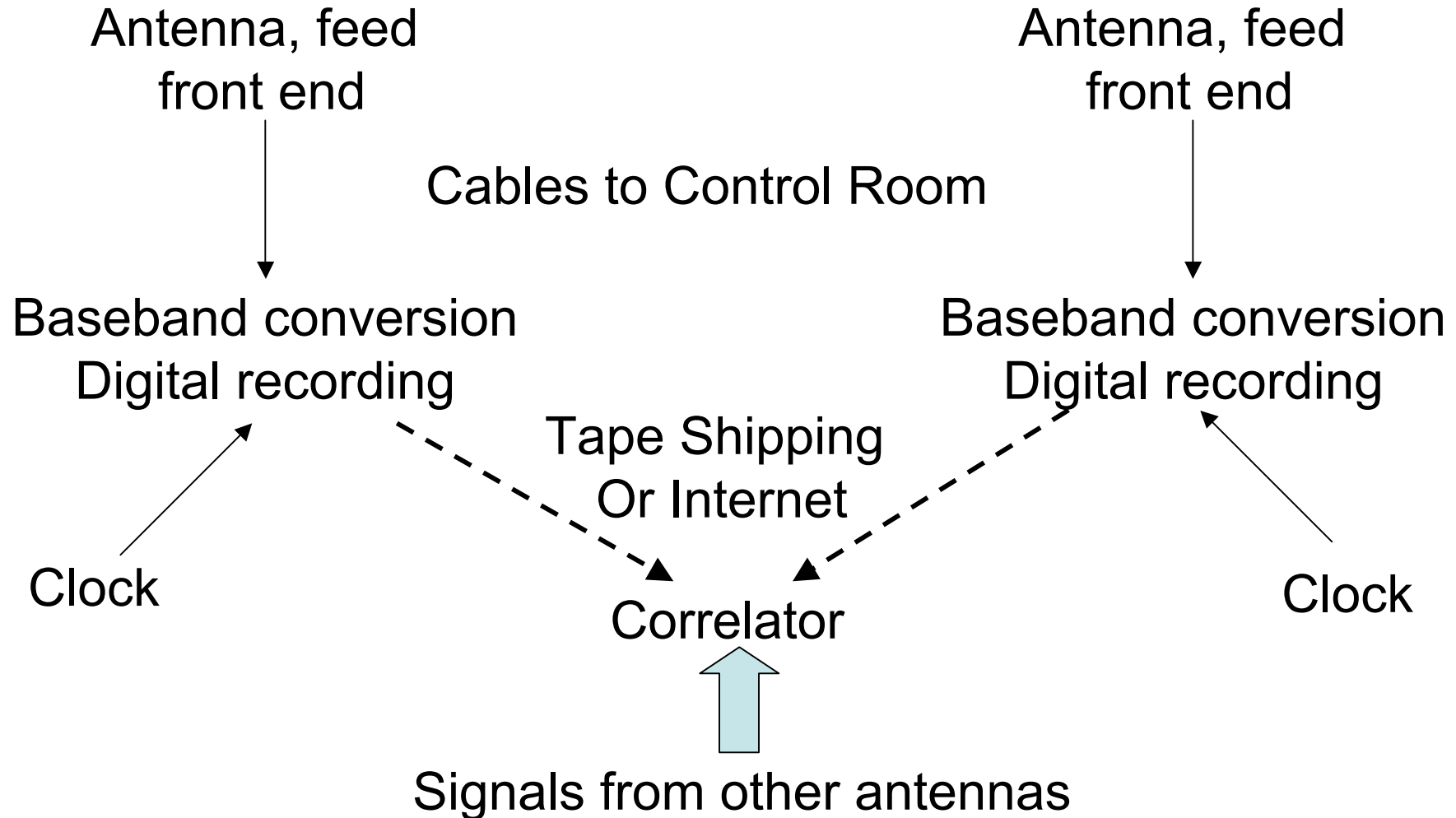
$$\tau = \frac{\vec{b} \cdot \hat{s}}{c} + \Delta\tau_c + \Delta\tau_I + \Delta\tau_{atm}$$



# Components of a VLBI system



# Basic VLBI System





# Antenna, Feed, Front End



- Antenna collects and focuses weak signals from Quasars
- Feed converts EM waves to electrical signals in wires.
- Front end amplifies the signal and reduces its frequency
- Photo: Antenna at Algonquin Park



# GGAO antenna at GSFC



# Composite antenna



- Approximately 15% of the weight of an equal size and stiffness Patriot antenna
- Kevlar and carbon fibre composites are about 10 times more thermally stable than steel and 30 times more thermally stable than aluminum.







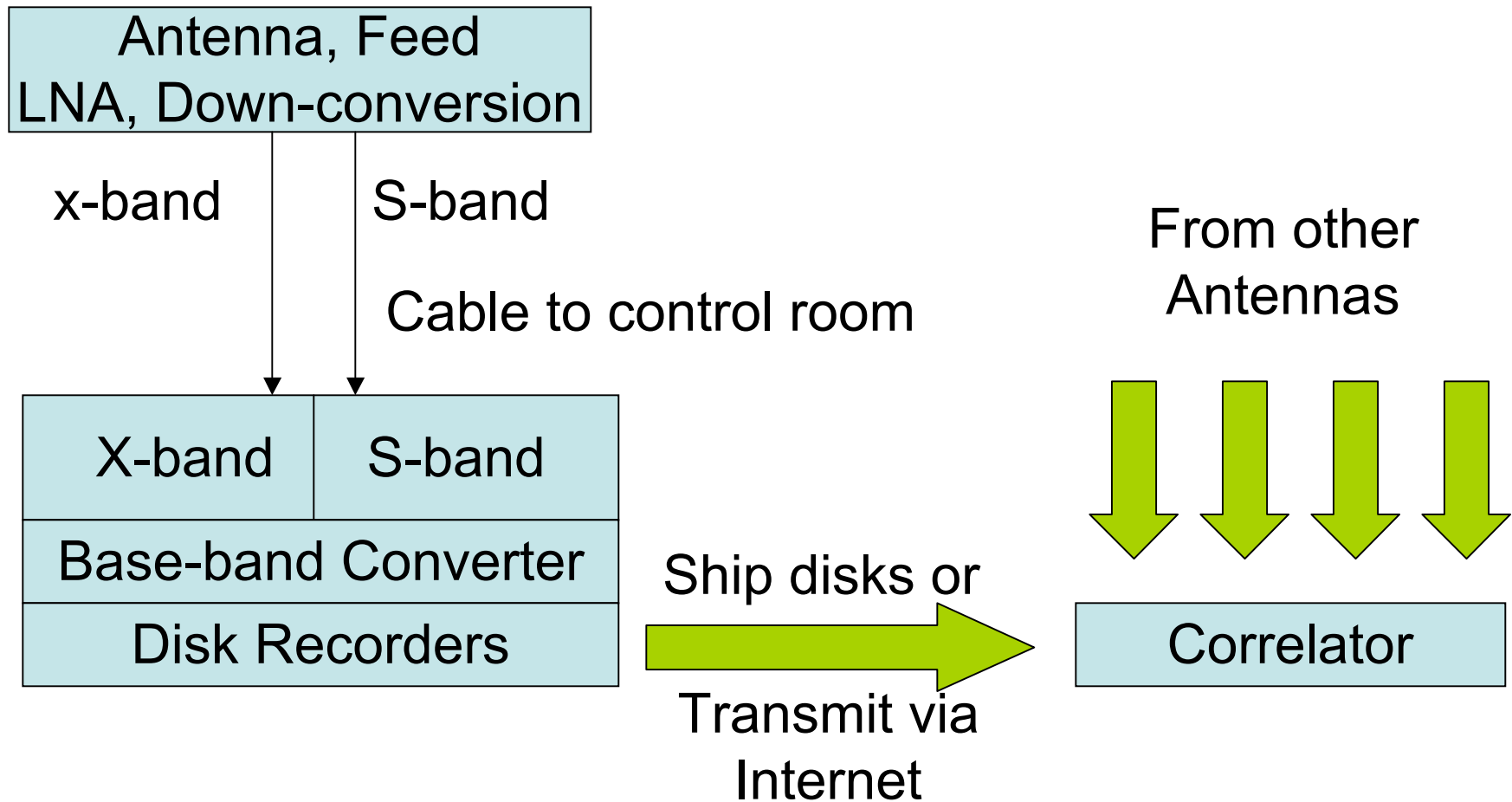




Canada's Natural P



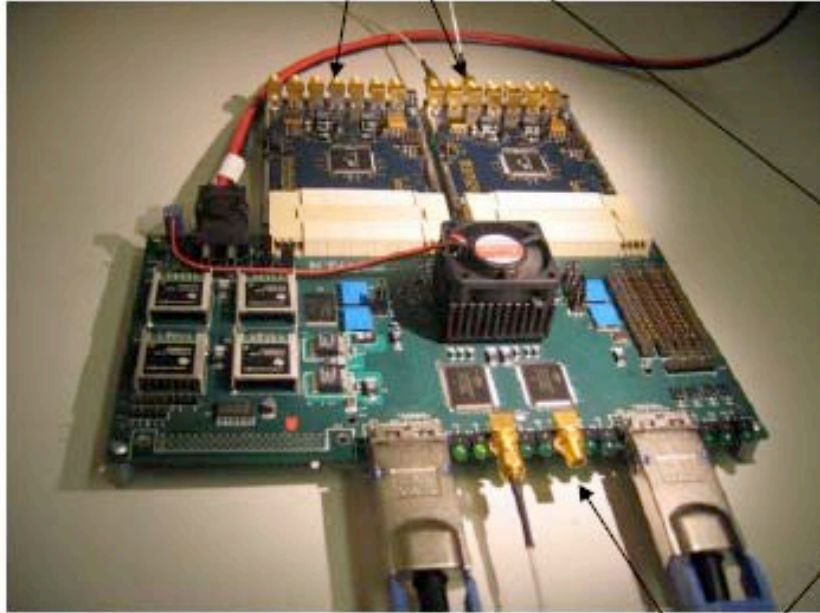
# Components of a VLBI system



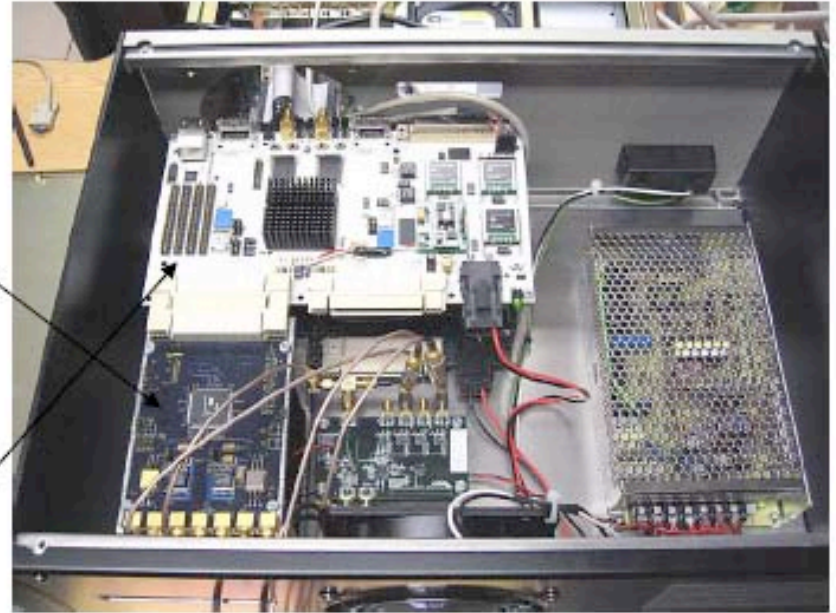
# Digital Back End – digitizer, baseband converter

## iBob/DBE1

### Sampler boards



iBOB



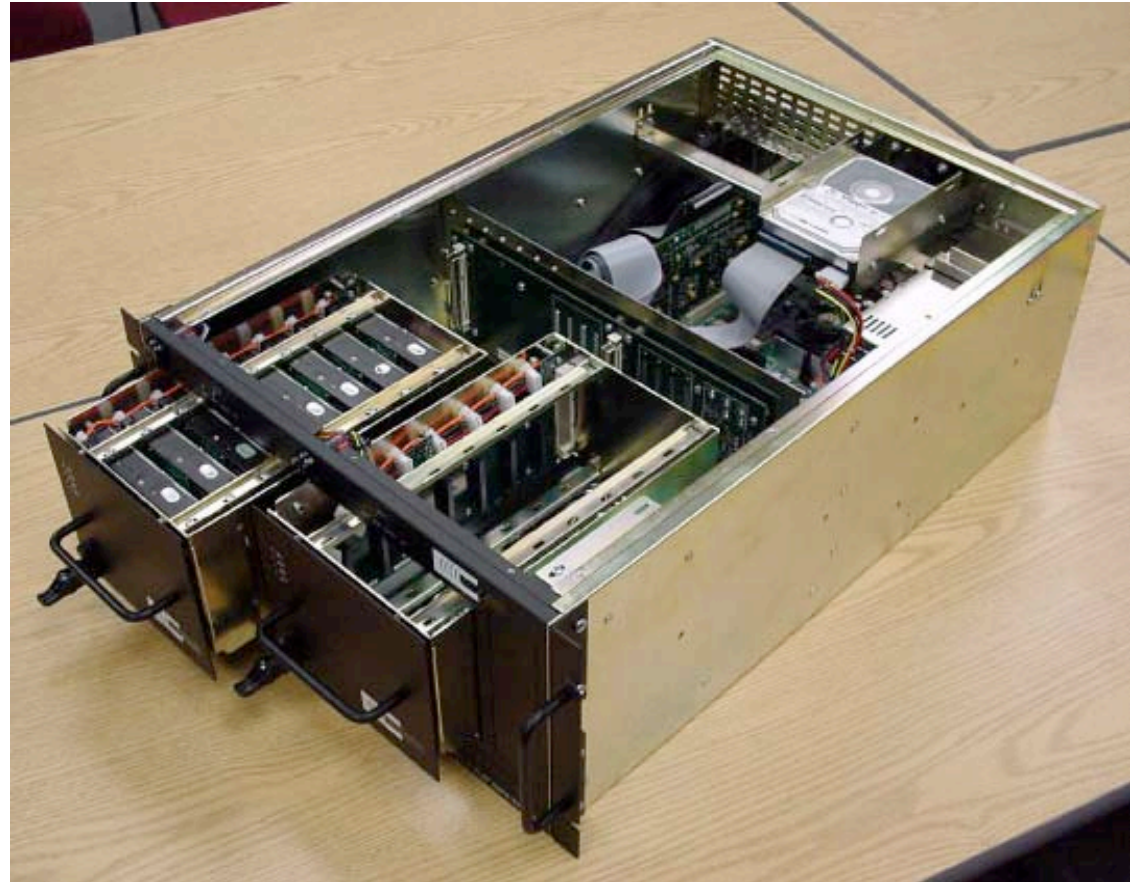
(two iBOB bds per chassis)



# Disk Record Unit – Mk5B



- Can record 1 Gbps data continuously for 24 h
- Mk5C is in design and will record at 4 Gbps
- Data rate has increased by nearly an order of magnitude per decade since 1967,

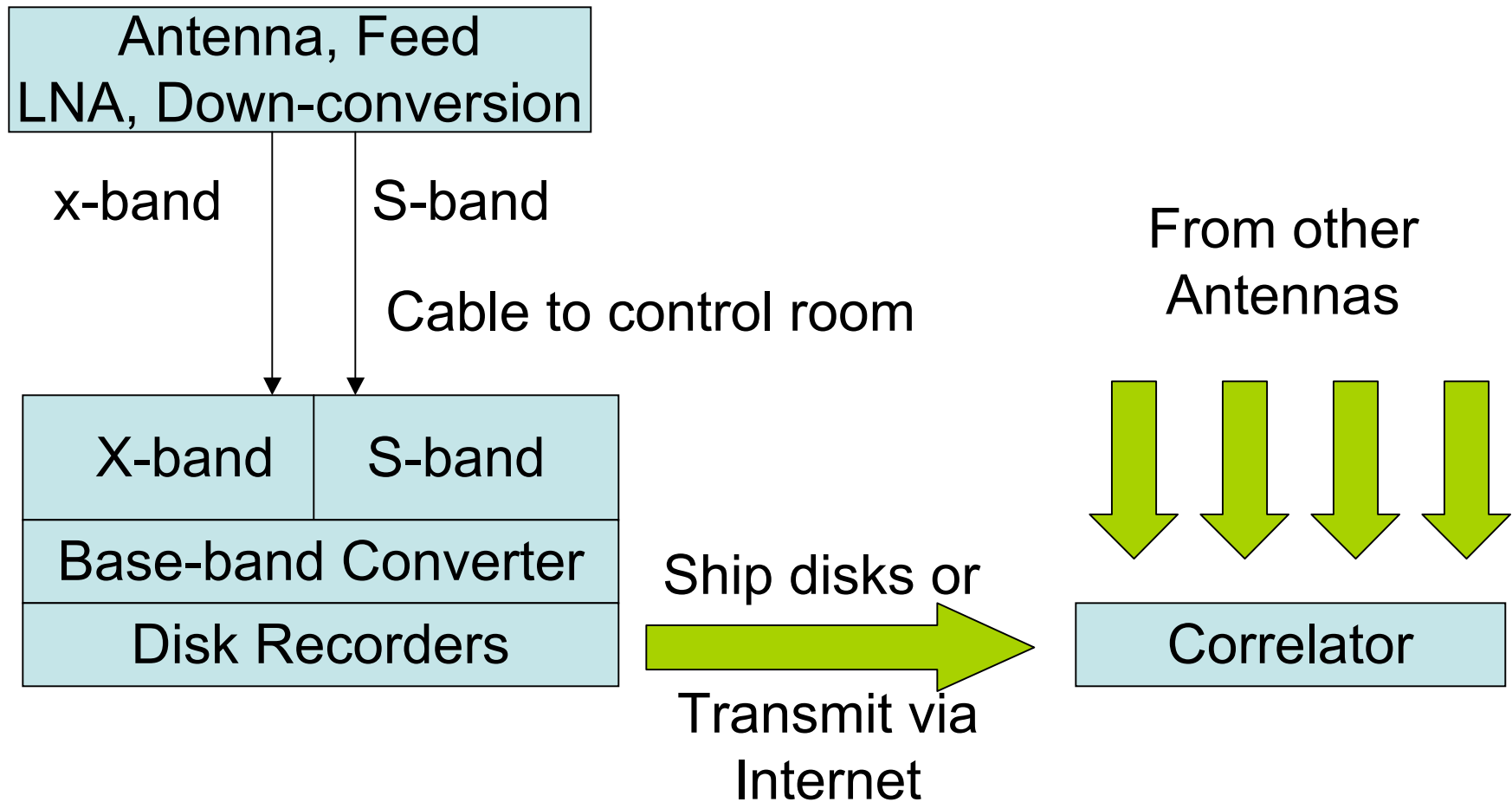


# eVLBI (Data Transmission by Internet)



- Required for quick turnaround to initial products
- Last km to antennas solved for many sites
- Sustained data rates near 1 Gbps achieved, but require vigilant monitoring of the light pathways
- 10 gige infrastructure expected to be widely available in the mid future -> achieves 8 Gbps VLBI2010 rates
- Risks
  - Cost and availability of research networks not known and definitely not guaranteed

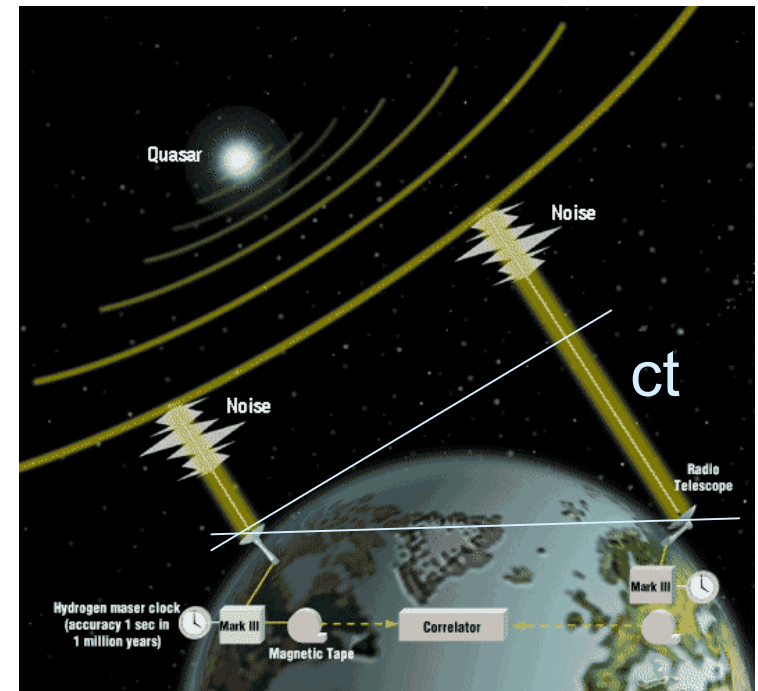
# Components of a VLBI system



# How does a VLBI correlator work?



- Signals are offset by a model delay to account for the interferometer geometry
- Signals are multiplied and accumulated (correlated) at different sample delays
- Positive correlation occurs when the two noise signals are aligned
- The offset of maximum correlation is added to the model delay to produce the output delay observable.





## Mk4 Correlator in Bonn Germany

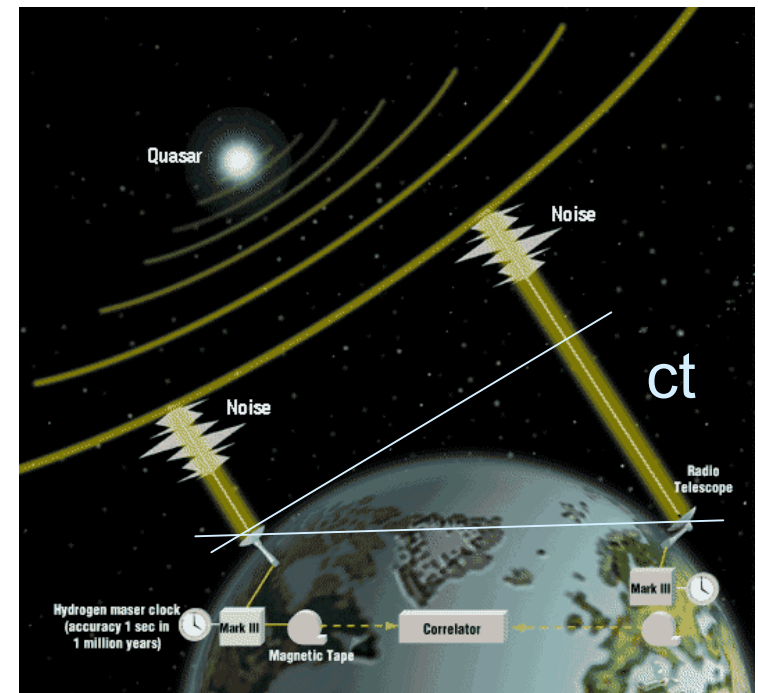


# The VLBI Delay Observable



$$\tau_{\vec{b}, \hat{s}}(t, f) = \frac{\vec{b}(t) \cdot \hat{s}}{c} + \Delta\tau_c + \Delta\tau_{Inst} + \Delta\tau_{atm} + \frac{K_{ion}}{f^2} + \Delta\tau_{\hat{s}}$$

- Major error sources are:
  - Neutral Atmosphere
  - Ionosphere
  - Instrumentation
  - Clocks
  - Source structure



# Error Source: Ionosphere

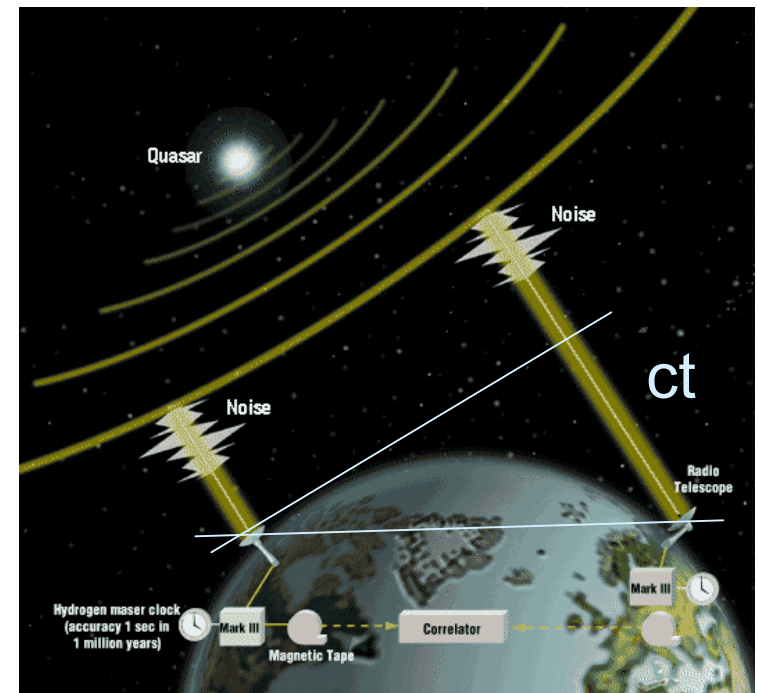


- Since the ionosphere is dispersive, i.e. changes with frequency, it can be calibrated by observing at two widely spaced frequencies, e.g. S-band (2.3 GHz) and X-band (8.5 GHz)
- $K$  is proportional to total line-of-sight electron content

$$\frac{K_{ion}}{f^2}$$

$$\hat{\tau}_S = \tau_{non} + \frac{K}{f_S^2} \quad \hat{\tau}_X = \tau_{non} + \frac{K}{f_X^2}$$

$$\tau_{non} = \frac{\hat{\tau}_X \cdot f_X^2 - \hat{\tau}_S \cdot f_S^2}{f_X^2 - f_S^2}$$



# Error Source: Neutral Atmosphere and Clocks



- Neutral Atmosphere has two components:

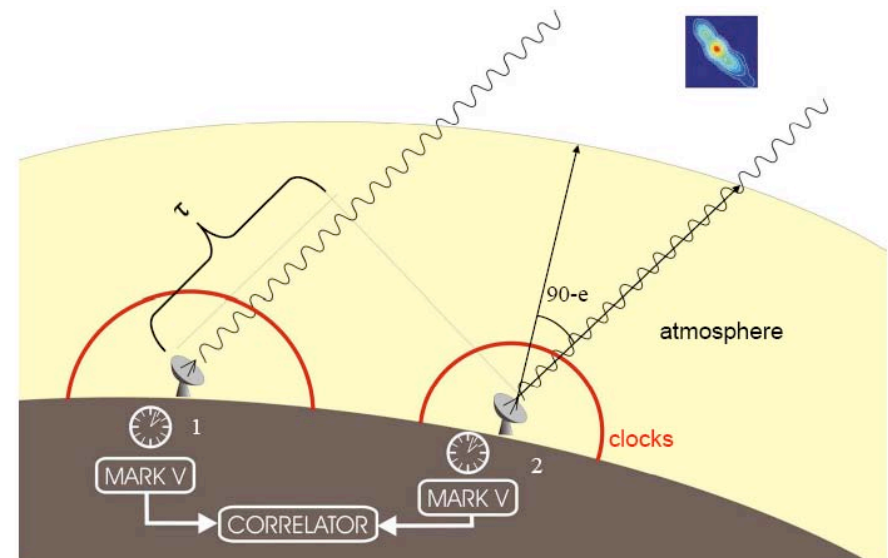
- Dry atmosphere
- Wet atmosphere

- The dry atmosphere is corrected using:

- Accurate dry mapping function
- The use of numerical weather data
- On-site pressure measurements

- The wet atmosphere
  - Water vapour radiometers

$$\Delta\tau_{atm}^D + \Delta\tau_{atm}^W$$



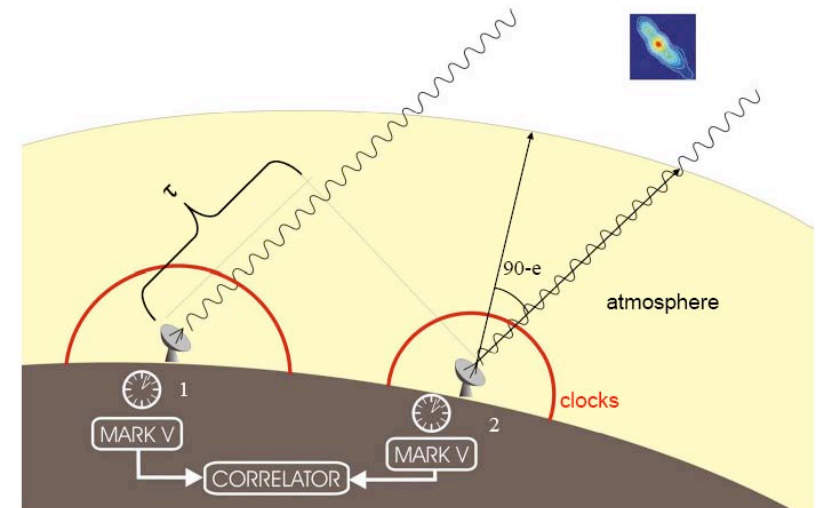


# Error Source: Neutral Atmosphere and Clocks



- The wet atmosphere is handled by estimating it from the VLBI delay data itself. This requires:
  - Accurate wet mapping functions
  - Clever scheduling to make the error sources separable, i.e.:
    - Geometry changes abruptly as the source changes
    - Atmosphere changes as elevation angle changes
    - Clocks change smoothly and continuously
    - Schedules switch rapidly between sources to vary geometry and elevation angle, making parameters separable.

$$\frac{\vec{b}(t) \cdot \hat{s}}{c} + \Delta\tau_c + \Delta\tau_{atm}^W$$

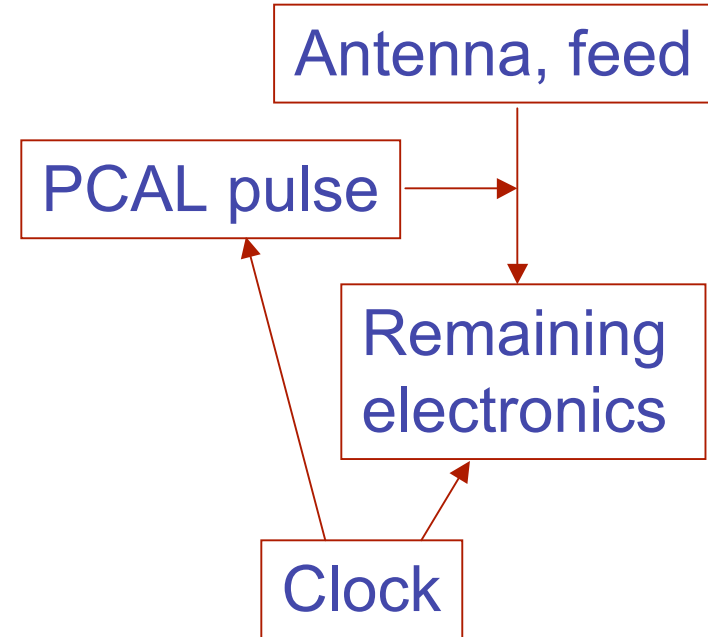


# Error Source: Instrumentation



- Instrumental errors are handled through phase calibration:
  - A very short pulse is injected early into the signal path
  - Since the pulse follows the same path as the signal, the delay of the signal through the electronics can be calibrated.
  - The limitations are:
    - It is only as stable as the calibration pulse
    - Not all of the system is calibrated

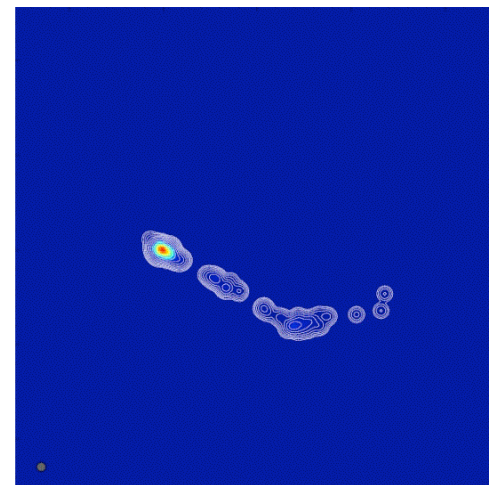
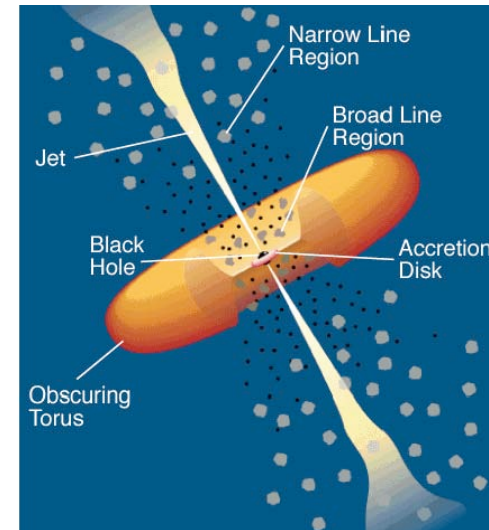
## Phase Calibration



# Error Source: Source Structure



- Sources are not in general points.
- To make matters worse, their structure changes with time.
- Lists have been carefully made of the best, most point-like and stable, sources
- If enough observations could be made, source structure could be corrected for.





- Thanks for being interested in VLBI and Geodesy!!!
- Questions?



- Please come to my talk tomorrow at 10:30 on the future directions of VLBI